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# Analysis of aberrations and performance evaluation of adaptive optics in two-photon light-sheet microscopy

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Abstract: Two-photon light-sheet microscopy (TP-LSM) system performa. A is greatly degraded by specimen-induced aberrations in illumination path, which limit the find of riew, axial resolution and excitation efficiency of the system. Adaptive optics (AO) is an effective method for attenuating these effects. For the design and evaluation of an AO system, a comprehensive analysis of the effects of aberrations is needed. In this paper, a TP-LSM system is simulated and riew indexes based on integral intensity are introduced for the evaluation of an aberrated lightharpoonup with the TP-LSM system are investigated with a numerical simulation method. Results show that high-oracle aberrations have little effect on the axial resolution and excitation efficiency of the system conly low-order components require correction. The random aberrations varied in strength with a depth of the specimens, so the number of corrected Zernike modes is variable. A general form, lair generated for the estimation of the number of modes that should be detected and corrected under different aberrations and different numerical aperture of the objective. The results can provide important guidance in the design and evaluation of AO units for TP-LSM systems.

Keywords: Adaptive optics; Aberration correction; Iwo-photon Light-sheet Microscopy

#### 1. Introduction

The optimization of biological microscopy system parameters, such as penetration depth, field of view (FOV), acquisition speed and photocing these cape fives in a single type of microscopic imaging system is currently challenging as optimizing these cape fives in a single type of microscopic imaging system is currently challenging as optimizing the end of these parameters may degrade the others. Two-photon light-sheet microscopy (TP-LSM) views in a single type of microscopic imaging system is currently challenging as optimizing the end of these parameters may degrade the others. Two-photon light-sheet microscopy (TP-LSM) views in a single type of microscopic imaging system is currently challenging as optimizing these parameters may degrade the others. Two-photon light-sheet microscopy (TP-LSM) views in a single type of microscopic imaging system is currently challenging as optimizing these cape and other photon in the advantages of different microscopy systems. TP-LSM systems use ultrafas near-infrared laser pulse to create a two-photon excitation light sheet, simultaneously achieving high imaging depth into biological tissues and high imaging speed with low phototoxicity and photocing high sheet, and photocing high sheet, simultaneously achieved and photocing high sheet.

A tradition of lights heet microscope illuminates a biological sample with a thin light sheet of visible light from the olde of the sample. The light sheet is imaged by a wide-field camera oriented orthogonally to the sheet. Given the orthogonal geometry of a light-sheet microscope, the whole light path contains two pasic parts: the illumination path and the detection path. Meanwhile, a TP-LSM system with trafast near infrared laser pulse instead of visible laser to create a two-photon excitation light sheet [1] and offers a submicron-scale axial resolution unlike a tradition light sheet microscope, which is limited to 2–8 µm axial resolution.

However, similar to other microscopy systems [3, 4], TP-LSM systems are affected by aberrations introduced by biological specimens and the two orthogonal light paths of the system are affected differently. Introducing adaptive optics (AO) to biological microscopy systems is one of the approaches

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for addressing such aberrations. Similar to AO technologies used in astronomical telescopes, specific wavefronts generated by wavefront correctors are used for offsetting decreased performance due to aberrations.

Over the past two decades, numerous adaptive optical microscopy systems have been implemented [5-8]. The AO technique has been widely used in two-photon microscopy [2-11] and light-sheet microscopy [12-14] for the past five years. Various detection and correction methods are applied to different systems, and many of them are of great value, such as fluorescopy guide star [15], scanning technique, and descanning technique [16]. However, in the field of biological microscopic imaging, biological specimens are multifarious, and the modes of construction and principles of existing systems greatly vary. Accordingly, the extent of performance degradation due to sample-induced aberrations vary among different systems.

The detection path of a TP-LSM system is the same as that of tractional 'ght-sheet microscope or any other plane imaging system. AO has been successfully applied to the illumination paths [17] and detection paths [12, 13] of traditional light-sheet microscopes. It were, how aberrations affect the illumination path of the TP-LSM is unclear because two-photon cocurs in this path and the quadratic dependence on the excitation laser intensity of co-photon-excited fluorescence make the influence of the aberrations on the illumination path marked different from that in the traditional one. In TP-LSM system, the illumination path is of vital in the excitation of the system. Knowledge about how abentations affect the illumination path is crucial to adaptive optical design. This knowledge can be used for determining the appropriate number of Zernike modes that needs correction, actuators and a corrector should contain, and sub-apertures that must be included by the Shack-Hartmann wavefro. It sensor (S-H WFS), which are the most important parameters of an adaptive optical system.

In this paper, we focus on aberrations in the illumination path of a TP-LSM system and analyze the effects of the aberrations by using amula, on method. Based on the point spread function (PSF) of the TP-LSM system, the two-photon  $n_s$  is simulated and two new evaluation indexes based on integral intensity are proposed. Then, we examine the effects of each individual Zernike mode of the aberrations and different strengths of random aberrations on the illumination path. The characteristics of the influence are clear, the number of Zernike modes that requires correction is determined, and a general formula is produced. These is sults provide guidance for adaptive optical design and aberration correction in TP-LSM systems.

#### 2. Methods

#### 2.1 Optical structure of P-L M.

The schematic diagram of fundamental TP-LSM system is shown in Fig. 1. In the illumination path, the ultrafa. 'near-in aread femtosecond pulse laser is emitted by a femtosecond laser source and is directed into the scan system employing axial (x-direction) and lateral (y-direction) scanning before entering the scan less (SL), tube lens (TL), and the rear pupil of the illumination objective (IO). In the detection path, the detection objective (DO), TL, and sCMOS camera are carefully aligned for the recording of the fluorescent signal excited by the illuminating laser.

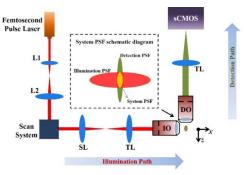


Fig. 1. Schematic diagram of the TP-LSM system.

In TP-LSM system, the virtual light sheet is generated by scanning in two directions and exposed by the sCMOS camera vertically. Compared with two-photon microsc py (TF^1), TP-LSM achieves larger FOV, lower photodamage, and higher acquisition speed without vausing obvious degradation in spatial resolution. However, the illumination FOV of a TP-LSM system cannot reach the desired requirement at high resolutions owing to the aberrations in the system cannot paths.

As shown in Fig. 2, aberrations are induced as light passes 'rough the specimen because of the variations in refractive index. In the illumination path, aborations are induced by region "a" in the specimen. By contrast, aberrations are induced by region "in the detection path. The two mutually perpendicular light paths suffer from different aberrations in different paths have varied influences on the postem. In TP-LSM, axial resolution results from the thickness of the light sheet, and lateral "esc auon is determined by detection optics [1]. Aberrations in the illumination path increases the "bickness of the light sheet and decreases the intensity, thus decreasing the axial resolution of the system and limiting the effective illumination FOV as illumination depth increases. Thus, the introduces of aberrations on the illumination path can be investigated quantitatively by evaluating thickness and intensity of the two-photon light sheet.

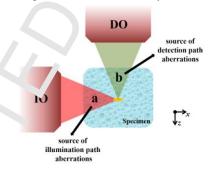


Fig. 2. Source of aberrations in different light path.

#### 2.2 Formation of two-p. ^to light sheet.

PSF must be calculated for the simulation of the two-photon light sheet. The system PSF of TP-LSM is calculated a follows:

$$PSF_{svs} = PSF_{ill} \times PSF_{det} \tag{1}$$

where  $\neg SF_{ill}$ , and  $PSF_{det}$  represent the system PSF, illumination PSF, and detection PSF respective. As shown in the inset of Fig. 1,  $PSF_{sys}$  is determined by the width of  $PSF_{ill}$  and  $PSF_{det}$  in the shown x-z plane. In this study, we solely focus on  $PSF_{ill}$ . The calculation method is provided below [18].

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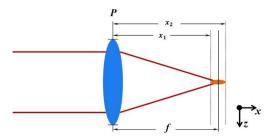


Fig. 3. The sketch map of light focused by a converging lens.

The parallel light is focused by objective lens of focal length f, as strown in Fig. 3. According to the Fresnel diffraction formula, the distribution  $U_x(y,z)$  near the back focal plane of the lens can be written as

$$U_{x}(y,z) = \frac{\exp[j\frac{k}{2x}(y^{2}+z^{2})]}{j\lambda x} \times \iint U_{l}(u,v) \exp[j\frac{k}{2x}(u^{2}+z^{2})] e^{-\frac{1}{2}} \left[-j\frac{2\pi}{\lambda x}(uy+vz)\right] du dv$$
 (2)

where x is the distance from the lens to the imaging plane. A containt phase factor has been dropped. Meanwhile,  $U_l(u,v)$  is the amplitude distribution behind the lens and x written as

$$U_{l} = U_{l}(u, v)P(u, v)e^{i\phi(u, v)} \exp\left[-\frac{1}{2}f(u^{2} + v^{2})\right]$$
(3)

where  $U_l(u,v)$  is the disturbance incident on the lens, v(u,v) is the aberrations in the pupil plane, and

P(u,v) is the pupil function. By changing the distance x from  $x_1$  to  $x_2$ , the amplitude distribution of the focus point along the x-axis can be obtained. The inclusive distribution can be written as

$$I_{x_1 \to x_2} = U_{x_1 \to x_2}(y, z) \times U_{x_1 \to x_2}(y, z)^*$$
(4)

For the two-photon excitation, the intentity of the detected fluorescence is

$$I_f = \kappa \delta_2 \eta I^2 \tag{5}$$

where  $\kappa$  is the collection efficancy of the imaging device,  $\delta_2$  is the two-photon cross-section, and  $\eta$  is the fluorescence quantum efficiency. Thus, the 3D effective illumination point spread function  $PSF_{iil}$  can be expressed as

$$P^{c}F_{ill}(x, y, z) = \kappa \delta_{2} \eta I^{2}_{x_{1} \to x_{2}} = \kappa \delta_{2} \eta [U_{x_{1} \to x_{2}}(y, z) \times U_{x_{1} \to x_{2}}(y, z)^{*}]^{2}$$
(6)

After the effect of the 'rection camera is considered, the intensity is integrated along the y-axis.

$$I_{camera}(x,z) = \int PSF_{ill}(x,y,z)dy$$
(7)

The virtual '.gut sheet is scanned along the *x*-axis, and the intensity distribution of the two-photon light sheet can be written as

$$I_{light-sheet}(z) = \int I_{camera}(x, z)dx = \iint PSF_{ill}(x, y, z)dydx$$
(8)

#### 2.3 Aberration description for TP-LSM.

In the adaptive optical system, the corrector is usually controlled with the slopes acquired by the S-H WFS. For convenience, the aberration decomposition is described by using the Zernike polynomial to express the aberrations. Therefore, the measured slopes should be transformed to the

coefficients of the Zernike polynomial. In the TP-LSM system, we use the same method to simulate the aberrations caused by specimens. The aberrations may be expressed as [19]

$$\phi(u,v) = \sum_{n=0}^{\infty} \sum_{m=0}^{n} a_{nm} Z_{n}^{m}(u,v) = \sum_{n=0}^{l} \sum_{m=0}^{n} a_{nm} Z_{n}^{m}(u,v) + \varepsilon$$
(9)

where  $Z_n^m$  is the orthonormal Zernike polynomial,  $\varepsilon$  is the residual error,  $a_{nm}$  is the corresponding coefficients, n and m are positive integers (including zero), and  $n-m \ge 0$  and ven. The index n represents the radial order, and m may be called the azimuthal order. By su' stitu  $m_{\rm b}$  Eq. (9) into Eq. (2) and Eq. (8), we can get the aberrated amplitude distribution of the illumination laser and the aberrated two-photon light sheet.

#### 3. Results

The influence of aberrations on the illumination path of the 'P-J M's investigated by performing a series of simulations. We simulate the formation of the two photon '.ght sheet on the basis of the integral effect of a camera and the Fourier transforming properties of objective lens. Assuming that the amplitude distributions of aberrations at different depths are uniform, we obtain the aberrated light sheet in the x-z plane. Our study focuses on the simulation and analysis of the aberrated light sheet, manifested in the effects of each Zernike mode and the random aberrations caused by biological specimens. All the simulations only apply to the plane that scan the beam in the lateral and axial direction.

In the simulation, the equivalent focal lengt. C the objective (f) is set to 5 mm, and the effective aperture (d) is changed. These steps ensure that different illumination numerical apertures  $(NA_{ill})$  are obtained. The wavelength of the femtosecond pulse laser is set to 1000 nm. The tip/tilt and defocus aberrations are removed from the wavefunctions in the simulations because they can only reposition the light sheet instead of distorting it.

#### 3.1 Simulation and evaluation inde . of two-poton light sheet.

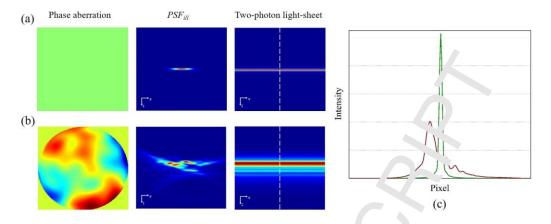


Fig. 4. The two-photon light-sheet simulation: (a) aberration-free phase,  $PSF_{iil}$ , and light sheet; (c) Intensity distribution of the two light enests  $a^{1}$  ong the z-axis.

In TP-LSM, the thickness of the light sheet determines the a. Treso ation of the system. With the thickening of the light sheet, the axial resolution decreases and a intensity is quadratically reduced, both leading to a limited FOV. To describe the influences of aberrations on axial resolution quantitatively, we define an evaluation index, which is the ratio between aberrated thickness and aberration-free thickness (*TR*) of the two-photon light of the two-p

$$TR = \frac{T}{1}$$
(10)

where  $T_A$  and  $T_0$  represent the thickness of aberra  $^{\circ}$ d and aberration-free light sheet, respectively. TR refers to the extent of the impact caused by approximation. The larger the TR is, the worse the effect of the aberration is. When the TR reaches 1, the aberration has no effect on the thickness of the light sheet and on the axial resolution of the TP-LSM ystem.

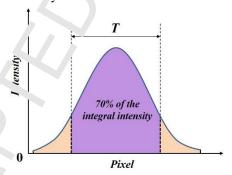


Fig. 5. Definition of the thickness  $T_{\bullet}$ 

It's import at to note that the full width at half maximum (FWHM) of the intensity distribution inaccurately discribes the thickness of the aberrated light sheet because the distribution curve of int analy may be distorted. Thus, we define a new parameter T as the thickness of light sheet. The parameter T is ne minimum width at 70% integral intensity, as shown in Fig. 5, and is equal to the FWHM of the two-photon light sheet in an aberration-free system. It can indicate the thickness of the aberrate.' light sheet in biological specimens accurately.

Simila 'y, the intensity of the light sheet indicates the excitation efficiency. To describe the influences quantitatively, we define another evaluation index: the ratio of aberrated intensity to aberration-free intensity (*IR*) of the two-photon light sheet, written as

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$$IR = \frac{I_A}{I_0} \tag{11}$$

where  $I_A$  and  $I_0$  represent the integral intensity of aberrated and aberration-free light. These along the z-axis, respectively. Similar to the Strehl ratio, the smaller the IR is, the worse the effect of the aberration is. When the IR reaches 1, the aberration has no effect on the excitation afficiency of the TP-LSM system.

## 3.2 Effects of each Zernike mode.

Considered the two-photon absorption effect, the first 27 Zernike makes and their distortion effects along x-axis are calculated and shown in Fig. 6. In this section, the ann-litude of aberration is appropriate to evaluate the aberration as it is specific and just contains one modes.

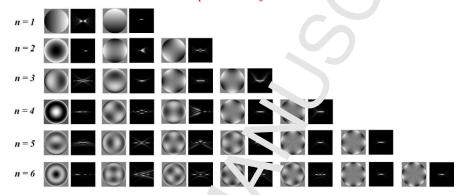


Fig. 6. The first 27 Zernike modes. For each mode, the distor  $\alpha$   $\alpha$  - $\alpha$ -proton  $PSF_{iil}$  along the propagating direction is shown. All results are calculated for amplitude  $\alpha$   $\alpha$   $\alpha$   $\alpha$  is the radial frequency of the aberration.

In the distorted  $PSF_{ill}$ , different Zernike mod's have varied effects on axial intensity distribution. That is, different Zernike modes play d'c' and roles in the thickness and intensity of two-photon light sheet. To quantificationally analyze the differences, we investigate the effects of the first 104 modes. For each mode, a series of amplitudes from  $\Omega$ ,  $\lambda$  to 2.0  $\lambda$  are investigated.

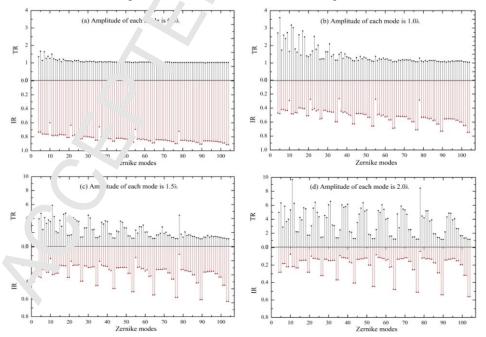


Fig. 7. Relationship between TR/IR and each Zernike mode. Amplitude of each mode is (a)  $0.5\lambda$ ; (b)  $1.0\lambda$ ; (c)  $1.5\lambda$ ; (d)  $2.0\lambda$ . Four representative results of the effects of different amplitudes on TR and IR are shown in Fig. 7.

First, we take the TR into account. In each Zernike mode, the greater the amplitude of the aberration is, the higher the TR is. Under relatively small amplitude (< 1.0  $\lambda$ ), the TR of low-order aberration is greater than that of the high-order aberration, and the high-order TR increases dramatically as the amplitude changes from 1.0  $\lambda$  to 2.0  $\lambda$  and keeps in line with the low-order TR. That is, the thickness of the two-photon light sheet is more sensitive to the low-order aberrations than to the high-order ones under small amplitudes. What's more, as indicated in Fig. 7(d), the aberrations effect the thickness of the light sheet regularly. In the same radial order n, the larger the azimuthal order n is, the smaller the TR is. When the value of m reaches n (high-order astigmatisms), the TR becomes n amplitudes and n is the smaller the n is n in the same radial order n, the larger the azimuthal order n is n in the smaller the n is n in the same radial order astigmatisms), the n is n in the smaller the n in the same radial order astigmatisms), the n in the smaller than the n in the same radial order n, the larger the azimuthal order n is n in the smaller than the n in the same radial order n is n in the smaller than the n in the n in the smaller than the n in the n in the smaller than

By considering the influence on intensity, we can easily understand the inverse relationship between TR and IR. As TR become larger, IR becomes smaller and vice velocity. When the two-photon light sheet becomes thicker, the dispersion of laser energy is in a relatively large area and the excitation efficiency will be reduced dramatically. However, as the figure indicates, an exception exists. The high-order spherical aberrations have little influence on the thickness of the two-photon light sheet even at a large amplitude but reduce the two-photon fluoresce. Line sity and degrade the image quality.

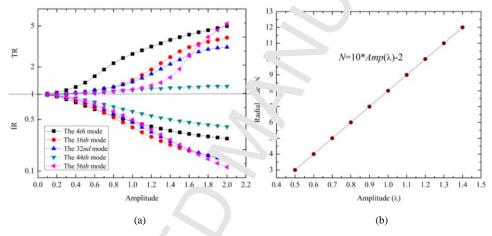


Fig. 8. (a) Representatives of different mc. es; ( $^{\dagger}$ ) Relationship between the amplitude Amp of each mode and the radial order N of aberration that can  $_{\circ}$   $_{\circ}$  rore  $_{\circ}$  rical to improve imaging performance in the correction.

Basing on the analyses bove, we conclude that the two-photon absorption effect can natively correct small-scale high-order (h. h radial frequency) aberrations. The importance of distorted illumination light is deer asec by the quadratic dependence of two-photon excited fluorescence on the excitation light intensity as a orophore excitation is spatially confined to only the highest intensity part of the beam, thus preserting axial resolution and fluorescence intensity even when the illumination light is distorted by as rations. To understand the conclusion in a more precise manner, we summarize the simulations of each Zernike mode. Several representatives of different modes are provided in Fig. 8(a). At increase d amplitude, different modes show different tendencies to change performance. We analyze the simulation and get the relationship between the amplitude *Amp* of each mode and the radial order *N* of observation that can be critical to the improvement of imaging performance in the correction. The result is fractor as a linear function, as shown in Fig. 8(b). Then, the effects of each Zernike mode are sum, ar zed in Table 1. In the analysis process, we think the aberration have a severe impact if TR reaches 1.5.

Table 1. Summary of the effects of aberrations on the thickness of two-photon light sheet

|                                  | Effect on thickness | Effect on intensity |
|----------------------------------|---------------------|---------------------|
| n-2m = 0 (spherical aberrations) | little effect       | great effect        |

| n = m (high-order astigmatisms)              | little effect | little effect |
|--|---------------|---------------|
| n≤N  |               |               |
| (1) 10 14 (1) 2 2 4 2 4 2                    | great effect  | great ef' sct |
| $(N=10*Amp(\lambda)-2; n-2m\neq 0; n\neq m)$ |               |               |

#### 3.3 Effects of random aberrations and the number of Zernike modes need to ! 2 co rected.

At every random aberrated wavefront, the amplitudes of high-order Zern. In modes are usually smaller than those of the low-order ones. In the field of biological microscopy, we amplitudes of 30–60 Zernike modes at the depth of 600  $\mu$ m are less than 0.5  $\lambda$  [10]. So only the worder aberrations should be corrected in TP-LSM. To determine the number of Zernike modes that should be corrected, we perform another simulation.

First, aberrations induced by biological specimens are simulated by using the Zernike polynomial. To make the simulations closer to the actual situation, we use the ab rration model in biological specimen proposed by M. Schwertner et al [20, 21]. The research incides that aberrations caused by the variations in biological specimen refractive index can be appro imate to random aberration. The Zernike mode standard deviation declines with rising order and big general behavior is found in all the specimens. In the adaptive optical microscopy system, we use RMS to represent the strength of the random aberrations induced by specimens at different dep bs as it represents the frequency character of the aberrated wavefronts. According to the results to the Your and by Kai Wang et al [9, 10] in the zebrafish embryos and mouse cortex, the aberration is weak an. RMS is approximately 0.1λ for the superficial specimen ( $<100 \mu m$ ;  $\lambda=1000 nm$  for the TP-LSM, Fe the deeper specimen, the aberration is moderate and RMS is approximately 0.5λ. When the immining position of the TP-LSM system goes deeper (>500 μm), the aberration is strong and RMS is proximately 1.0λ. Based on the aberration model above, 1000 wavefronts are generated runningly for each aberration strength, and the average TR and IR of the 1000 sets of wavefronts ar used to obtain statistical results. The closer to 1 the TR and IR values are, the better the correction effect of the method is. The relation among TR, IR, and the number of corrected Zernike modes J und or random aberration with variable strength is shown in Fig. 9.

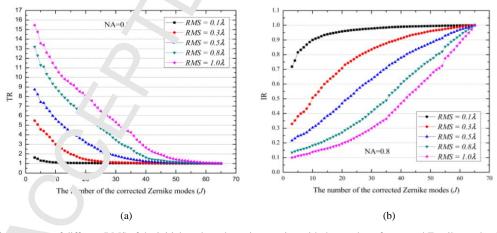


Fig.  $^{\circ}$ .  $^{\circ}$  different RMS of the initial random aberration varying with the number of corrected Zernike modes. NA of the illumination objective is 0.8.

As is shown in Fig. 9(a), TR and IR change with the number of corrected Zernike modes for varying RMS of the random aberration. Practically, it can be assumed that the correction of AO unit is ideal when TR is smaller than 1.1 and IR is larger than 0.8. For weak aberration, RMS=0.1 $\lambda$ , the first 11 modes need to be corrected to achieve ideal correction. Considering a stronger turbulence,

RMS=0.3λ, TR is 1.09 and IR is 0.86 with 33 Zernike modes corrected. However, for moderate or strong aberration, 49 or more Zernike modes should be corrected to make it.

In TP-LSM, NA<sub>ill</sub> is an important parameter. The effects of random aberrations vary among different NA<sub>ill</sub> values. Fig. 10(a) illustrates the meaning of NA dependent aberration. For the same objective, the working distance (WD) is fixed. When the illumination aperture decreases, the size of the light cone passing through the specimen decreases, and the system will suffer a more moderate aberration. To simulate this phenomenon, a single random aberration is generated to the high NA<sub>ill</sub> and then cropped in the simulation of a lower NA<sub>ill</sub>. The two-photon light sheets and the correction process are simulated under three different NA<sub>ill</sub> (Fig. 10(b-d)). The effect is sever, at high NA<sub>ill</sub>, and the same correction effect can be obtained by increasing the number Zernike modes for correction.

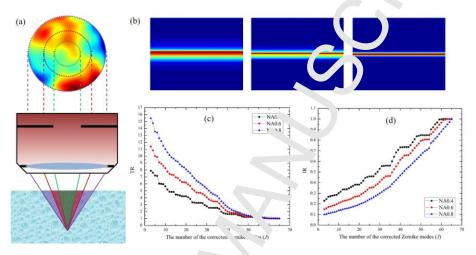


Fig. 10. (a) Illustration of NA dependent aberrations; (b) "vo-photon light-sheet under different NA<sub>ill</sub>; (c) TR and (d) IR varying with the number of corrected Zernike modes under different NA<sub>ill</sub>.

To investigate the influence more fully, we make a specific comparison of the correction process between different NA<sub>ill</sub> and different support of the random aberrations. The aberration for lower NA<sub>ill</sub> is cropped from the highest one, and the corresponding RMS is changed and calculated. The detailed number variation values of corrected Ze nike modes to keep TR below 1.1 and IR above 0.8 are provided in Fig. 11. To make the results more universal, we fit the points as power functions. The fitting equations are shown in the agure at the same time. With the three fitting equations, a general formula between the number of corrected Zernike modes (J), the RMS of the random aberrations, and the NA<sub>ill</sub> can be made and atten as

$$J = 58 \times \left[ 1 - \left| RMS(\lambda) - 1.03 \right|^{(2 \times NA_{H} + 1.2)} \right]$$
(12)

With the general formula above, the number of Zernike modes that the AO system must detect and correct under dufferent scales of random aberrations and different NA<sub>ill</sub> values can be estimated. This formula plates a great part in the design and performance evaluation of an AO system in TP-LSM.

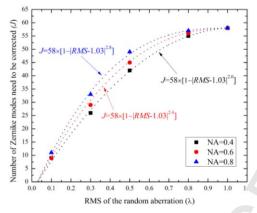


Fig. 11. Number of Zernike modes that requires correction under random aberration won variable strength and different NA of the objective.

#### 4. Discussion

In this paper, a two-photon light sheet is simulated, and the evaluation indexes of the light sheet are proposed. The influence of aberrations on the illumination part of the TP-LSM is studied. The effects of each Zernike modes and random aberrations on the two-photon light sheet are investigated, and the relationship among the amplitude of each mode, the Rock of the random aberration, the number of the corrected Zernike modes, the NA of illumination objective, the thickness and intensity of the light sheet is derived.

Based on the relationship above, a series of 1 alts can be produced. First, in TP-LSM, the high-order aberrations have little effect on the a. a. resolution of the system because of the quadratic dependence on the excitation laser intensity of a vo-photon-excited fluorescence. Thus, only the low-order components should be corrected. So and, as the aberrations at different depths of the biological specimens vary in strength, the number of Zernike modes that needs correction is variable. Third, when the NA<sub>ill</sub> gets large, the nickness of the light sheet becomes increasingly sensitive to the same aberration. This result indicates that small NA should be used in the illumination path of TP-LSM. Fourth, a general form that it mades for the estimation of the number of modes that should be corrected in different case. Ac ording to the formula, although varing with the strength of aberrations induced by the specimens and the objective, the number of Zernike modes that should be corrected can be fixed at the contribute to the design of wavefront sensor and corrector and the performance evaluation of Accinit for TP-LSM.

In the AO system, the sub-aperture number of S-H WFS determines the number of Zernike modes that can be measured, and the element number of corrector determines the number of Zernike modes that can be confected. It ence, the general formula in the TP-LSM can be used in the selection of a sub-aperture number for S-H WFS and element number for the wavefront corrector. For the application of AO on the illustration path of TP-LSM, a S-H WFS with 10×10 sub-apertures is sufficient to measure 58 Ternike modes, and a deformable mirror (DM) with more than 100 actuators or a liquid crystal path of the correction of aberrations.

On use basis of the conclusions above, an experimental TP-LSM system with an AO unit will be built in the future work. By applying the AO technique, a fluorescence microscopy system can achieve large FOV and high spatial resolution.

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