

Spherical mirror testing by phase retrieval wavefront sensor

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ABSTRACT

In order to verify the estimated wavefront ability of the phase retrieval wavefront sensor (PRWS), a measured spherical mirror of experiment platform was set up with the method of PRWS. PRWS technology is based on the focal plane image information wavefront solver in the focal plane wavefront measured technology, whose principle is sampling a number of the given defocus images; get the wavefront phase information by solving the optical system wavefront with Fourier optical diffractive theory and mathematics optimization. In order to validate the veracity of PRWS, both the PRWS measurement results and ZYGO interferometer measurement results were compared; experimental results demonstrate that agreement is obtained among the errors distribution, PV value and RMS value of ZYGO interferometer, so PRWS technology can effectively estimate the aberrations of spherical mirror.

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1. Introduction

In the machining processing of large-scale optics mirror, in situ real-time estimation and alignment with the use of the optical system during dynamic measurement of wave aberrations are difficult to accomplish for the present tradition optical inspection equipment [1–3]. In order to control the optical quality of the telescope, we need a simple and with high-accuracy method. This article proposed phase retrieval wavefront sensor (PRWS) technology [4–8] based on the focal plane image information wavefront solver in the focal plane estimated wavefront technology, whose principle is given by sampling a number of the defocus images [9–14]; solve the optical system wavefront by Fourier optical methods. System hardware is simple, free from the environment (especially the vibration) influence of optical components and systems for dynamic estimation, real-time display of measurements [15], which has good application prospects in the field of the optical processing, system alignment, active optics, adapt optics and others.

In order to verify the estimated wavefront ability of PRWS, this paper set up a measured spherical mirror of experiment platform with the method of PRWS based on phase retrieval (PR) theory research and experimental verification. This paper compared PRWS measurement results with ZYGO interferometer [16–22] measurement results, experimental results demonstrate that agreement is obtained among the errors distribution, PV value and RMS value of ZYGO interferometer, so using PRWS technology can effectively

test the spherical mirror aberration, which illustrates the feasibility and accuracy of PRWS measurement methods.

This paper is organized as follows: the theory of PRWS is presented in Section 2, the design of experiment in Section 3 and the summary in Section 4.

2. The principles of PRWS

Phase retrieval (PR) system is the wavefront detector of a focal plane waves; a laser spot light on the object plane is a target designated from the focal plane image acquisition, use the acquired image, the defocus of the corresponding image, known pupil size and shape to reverse solve the aberration of the optical system [23]. The structure of the PR system is shown in Fig. 1.

Assuming that the aperture of a measured optical system is D , the focal length is Z , the center wavelength of the laser light source is λ , whose pupil constraint function is $|f(x)|$, the generalized pupil function for focus plane is

$$f(x) = |f(x)| \exp[i\theta(x)], \quad (1)$$

where θ is wavefront distortion and can be obtained with Zernike polynomial fitting: $\theta(x) = \sum_n \alpha_n Z_n(x)$, the real number α_n represents the first n terms of polynomial coefficients, Z_n indicates the first n terms of Zernike polynomials basement.

For linear optical system, when the generalized pupil $f(x)$ whose defocus is δ in the plane, the impulse response function $F(u)$ is

$$F(u) = |F(u)| \exp[i\psi(u)] = \mathcal{F} \{f(x) \exp[i\theta(x, \delta)]\}, \quad (2)$$

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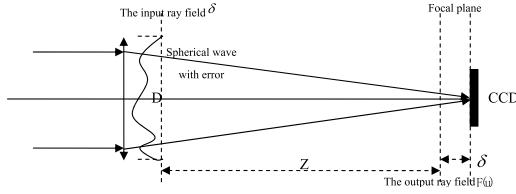


Fig. 1. Schematic of optical path of PR.

where x is the coordinates of the pupil domain, u is the coordinates of the image domain, both of them are two-dimensional vector field coordinates. ψ is the phase part of the impulse response, \mathcal{F} is two-dimensional Fourier transform, $\varepsilon(x, \delta)$ is wavefront aberration caused by defocus δ in the position x .

For a PR system, $|f(x)|$ of Eq. (1) is the priori conditions of a known optical system, corresponds to the size and shape of the pupil. $|F(u)|^2$ is the image collected by CCD where the defocus is δ . Therefore, the purpose that we estimate wavefront by PR is to get α_n by the above known quantity. So formal description of the problem for: $|f(x)|, \delta_1, |F_1(u)|^2, \delta_2, |F_2(u)|^2, \dots, \delta_M, |F_M(u)|^2$ are known. Image acquisition distance from the focal plane at $\delta_1, \delta_2, \dots, \delta_M$, respectively, is $|F_1(u)|^2, |F_2(u)|^2, \dots, |F_M(u)|^2$.

The objective function and the partial derivative of PR objective function with respect to α_n , respectively, is formulas (3) and (4)

$$B_k = E_{Fk}^2 = N^{-2} \sum_{m=1}^M \sum_u [|G_{m,k}(u)| - |F(u)|]^2, \quad (3)$$

$$\partial_{\alpha_n} B_k = -2 \sum_m \sum_x |f(x)| |g'_{m,k}(x)| \sin[\theta'_{m,k}(x) - \theta_{m,k}(x)] Z_n(x). \quad (4)$$

With the objective function (3) and its impact on the Zernike coefficient derivative (4), we can use the mathematical optimization algorithm to solve various Zernike wavefront coefficient values, here we use LBFGS algorithm that the phase diversity (PD) [24–30] experiment has been able to solve.

3. The design of the experiments

3.1. Experimental theory and components

The schematic diagram of PRWS is shown in Fig. 2. Gaussian beam emitted from the laser through the pinhole into a spherical wave, passes through the lens 2 into a parallel light, the light projects in the prism through the prism is divided into two parts, a part need not be considered, another part of the parallel reflects after a light through the lens 1 converge after the measured mirror, the reflected beam with phase [31–35] information (i.e. aberration), again divided into two groups by a beam splitter, one part back-track, the other part through the converging lens 3 converged at CCD camera, which is placed on a movable platform, move along the optical axis and the angle of the camera posture fine-tuning to get the focus before and after receiving a different amount of defocus

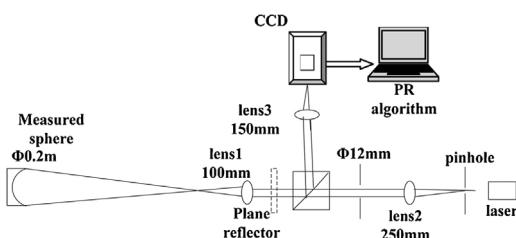


Fig. 2. Schematic diagram of PRWS.

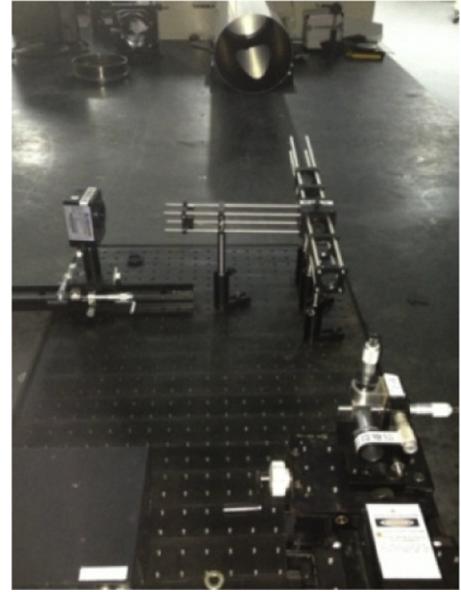


Fig. 3. The experimental system of PRWS.

images [36–39], used in realizing the estimated wavefront based on the PR. We can obtain aberration of the measured spherical mirror with the PR algorithm.

The focal length of measured spherical mirror is 0.2 m, the center wavelength is 632.5 nm, focal length of 3 in the experimental system is 0.15 m, the exit pupil caliber is 0.012 m, and the depth of focus is about 0.286 mm. In the experiment, the defocus we select is $0, \pm 1, \pm 1.5, \pm 2$ mm. Camera pixel size is $6.45 \mu\text{m}$, each defocus position, respectively, intercept 128×128 pixel size of target region, the exposure time is 20 ms, the accuracy of mobile platform is $\pm 5 \mu\text{m}$. The experimental optical path is shown in Fig. 3.

3.2. Experimental procedures

In the process of the entire experiment, we not only prove the estimation ability of PRWS, but also prove the accuracy of PRWS, therefore during the experimental design, in order to ensure that the position of the measured spherical mirror during the whole experiment is invariable, we need to find the good distance between spherical mirror with PRWS devices and between spherical mirror with ZYGO, respectively, based on the focal lengths of the measured spherical mirror. Then separately measure the spherical mirror with PRWS and ZYGO.

Step1: build experiment system. Firstly, fix the 0.2 m telescope as shown in Fig. 4; secondly, place the laser and the position of the pinhole as shown in Fig. 5. The requirements of laser position are: laser light and the center of the telescope mirror are coaxial. After regulating, put the microscope objective and a pinhole placed in front of the laser making the light from the pinhole in an ideal spherical wave. Lastly, reconfirm the laser, pinhole, spherical mirror are coaxial, as shown in Fig. 6.



Fig. 4. Tested telescope.

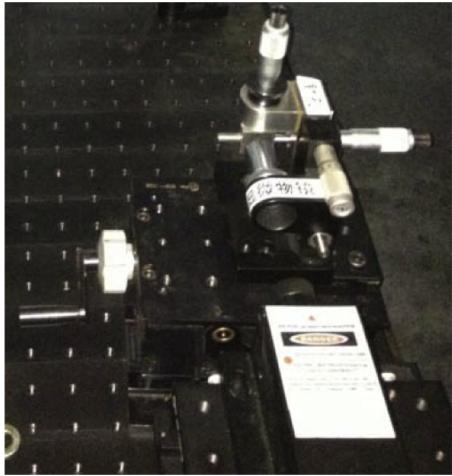


Fig. 5. laser and pinhole.

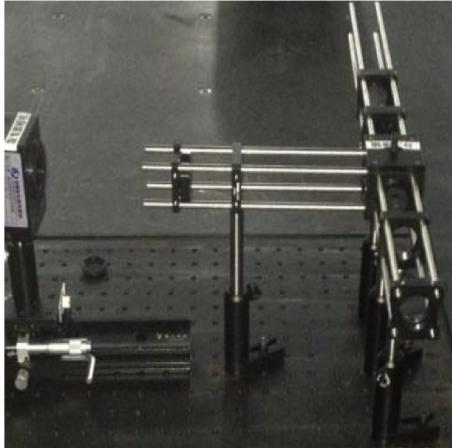


Fig. 6. Coaxial system.

Step 2: add the lens 1, and adjust the position of the lens 1 to make it to be coaxial with the laser and pinhole until the spherical wave is emitted from the pinhole through the lens 1 into parallel light;

Step 3: add the splitting prism, light bar and the lens 2 in the optical path, and make them be coaxial with the measured spherical mirror and pinhole, then according to the parallelism of the return light beam from splitter to adjust the position of the lens 2;

Step 4: adjust the position of the lens 3 and the camera to make the light beam reflect from the measured spherical light through the prism to enter into the camera;

Step 5: after adjusting the system, adjust the translation stage to make the point of light least, then record the location and use it at the position of the focal point;

Step 6: capture images with the camera at the position of the defocus distances of $\pm 1, \pm 1.5, \pm 2$ mm;

Step 7: put the plane mirror between the lens 1 and the prism, repeat the above steps to capture images with the camera;

Step 8: dispose the images collected of the sixth and seventh steps with the PD algorithm, get the aberration of the measured spherical mirror;

Step 9: ensure that the position of the measured spherical mirror remains unchanged, use the ZYGO interferometer estimate the measured spherical mirror. **Fig. 7** is a diagram of the estimation experiment with the ZYGO interferometer.



Fig. 7. The experimental system of ZYGO interferometer measurement.

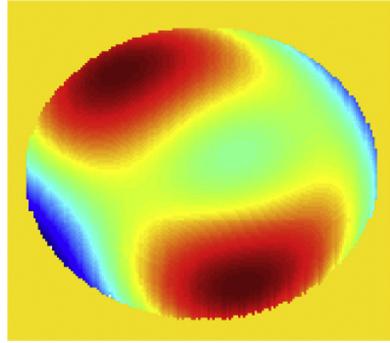


Fig. 8. Result of PRWS, $RMS = 0.272\lambda$, $PV = 1.608\lambda$.

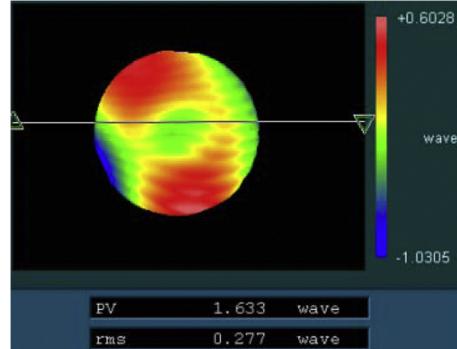


Fig. 9. Result of ZYGO interferometer, $RMS = 0.277\lambda$, $PV = 1.633\lambda$.

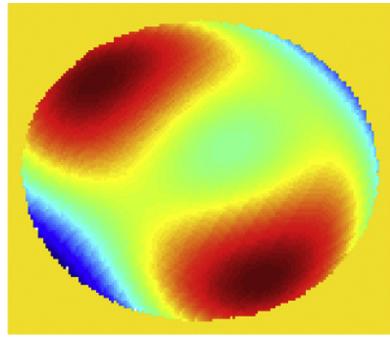


Fig. 10. Result of PRWS after rotation.

3.3. Experimental results and discussion

We dispose the collected seven images with the PR algorithm, obtain the measured result of the spherical mirror as shown in **Fig. 8**. The measured result with the ZYGO interferometer is shown in **Fig. 9**.

In order to illustrate the accuracy and viability of the PRWS better, we rotated the spherical mirror a definite degree, and then estimate the spherical mirror with steps one to nine, the obtained measurement results are shown in **Figs. 10 and 11**.

In order to illustrate the accuracy and repeatability of PR, we rotated the spherical mirror some degree, and then estimate the

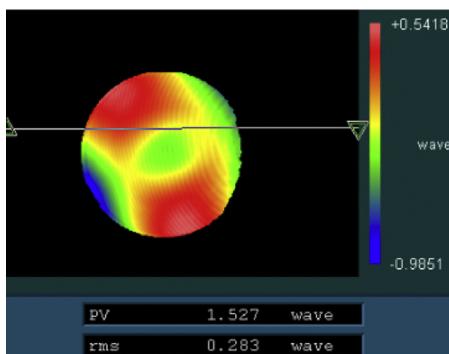


Fig. 11. Result of ZYGO interferometer after rotation, RMS = 0.280λ , PV = 1.501λ , RMS = 0.283λ , PV = 1.527λ .

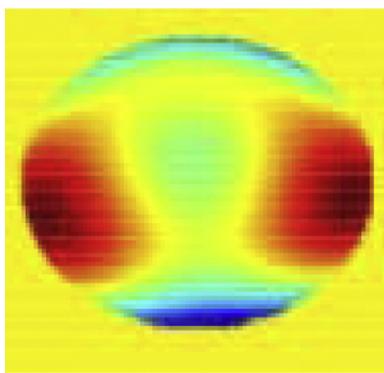


Fig. 12. Result of PR measurement after rotation, RMS = 0.271λ , PV = 1.659λ .

Table 1
Comparison of Zernike polynomials.

ITEMS	Zernike	ZYGO	PRWS
1	Piston	0	0
2	X tilt	0	0
3	Y tilt	0	0
4	Defocus	0	0
5	Ast x	-0.380	-0.372
6	Ast Y	-0.552	-0.543
7	Coma x	0.018	-0.017
8	Coma Y	0.044	0.043
9	Primary Spherical	-0.217	-0.213
10	Trefoil x	-0.013	-0.012
11	Trefoil y	0.240	0.235
12	Secondary Ast x	0.045	0.044
13	Secondary Ast y	0.028	0.027
14	Secondary coma x	-0.065	-0.063
15	Secondary coma y	-0.007	-0.006
16	Secondary Spherical	0.100	0.099
17	Tetrafoil x	0.030	0.029
18	Tetrafoil y	0.014	0.013
19	Secondary Trefoil x	-0.020	-0.019

spherical mirror with steps one to nine, the obtained measurement results are shown in Fig. 12.

In order to verify the accuracy of PRWS we compared the 19th Zernike polynomials of Figs. 10 and 11 and get the results as shown in Table 1.

From the results of Table 1, we can see that the Zernike polynomials of PRWS and ZYGO are linear relationship, which is the same as the difference RMS between PRWS and ZYGO. From the measurement result after rotating, as shown in Fig. 12, the results of estimated wavefront before and after rotation are in agreement, which verifies the repeatability and effectiveness of PRWS measurement. Seeing from Figs. 8–11, there is an agreement among the errors distribution, PV value and RMS value of ZYGO

interferometer, which explains the feasibility and accuracy of the PRWS measurement methods.

4. Conclusions

This paper set up an estimation spherical mirror of experiment platform with the method of PRWS, from the contrast of the measurement results before and after rotation, which verified the PRWS measurement repeatability and effectiveness. In order to validate the veracity of PRWS, this paper compared PRWS measurement results with ZYGO interferometer measurement results, experimental results demonstrate that agreement is obtained among the errors distribution, PV value and RMS value of ZYGO interferometer, so using PRWS technology can effectively estimate the spherical mirror aberration, which explains the feasibility and accuracy of the PRWS measurement methods, which provides the feasibility to data support for our later search.

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